

NIF



Nuclear-Plasma Interactions at NIF and accelerator labs

**Exotic Beams Summer School
Lawrence Berkley National Laboratory
Berkeley CA**

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LLNL PRES pending

Lawrence Livermore National Laboratory • National Ignition Facility & Photon Science

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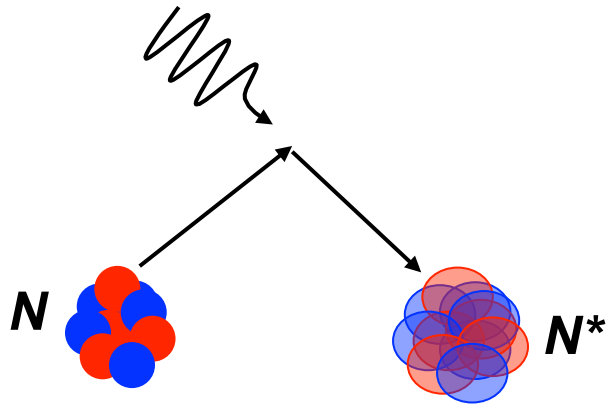
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Introduction

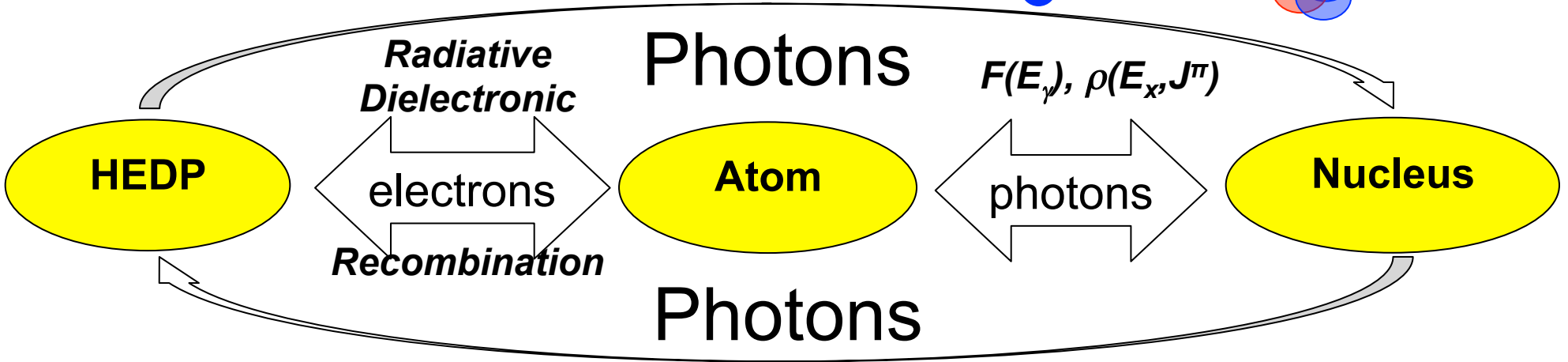
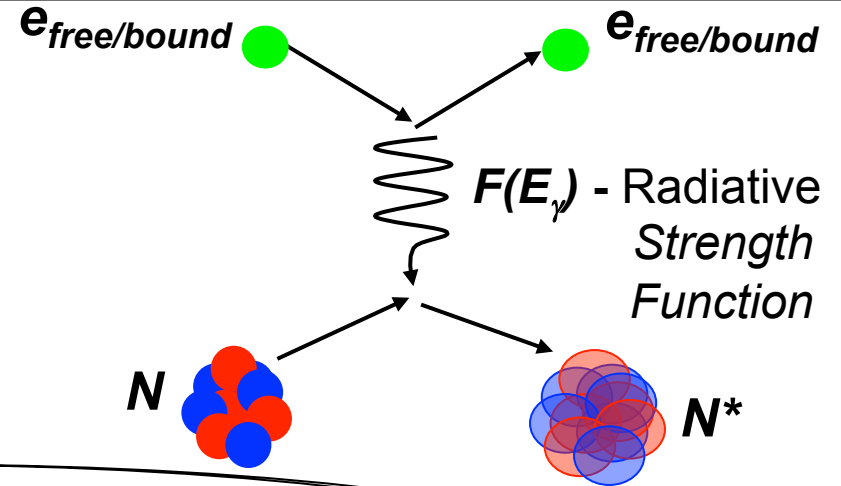
- Nuclear-plasma interactions in Neutron-rich High Energy Density Plasmas (nHEDP)
- Nucleosynthesis in stellar nHEDPs
- Nucleosynthesis at the National Ignition Facility
- Results from NIF – $^{196m}\text{Au}/^{196g}\text{Au}$
- Other planned and potential experiments
 - NIF-based exploding pusher with ^{134}Xe
 - Accelerator-based using Au beams
 - *Petawatt-laser beam-target experiment (Au)*
- ***Nuclear Level Density and Radiative Strength is crucial to understanding the formation of elements in nHEDPs***

Nuclear-Plasma Interactions (NPI) can excite nuclear states with energies comparable to those of the surrounding plasma

Photo-absorption
Time Reverse: γ -ray decay

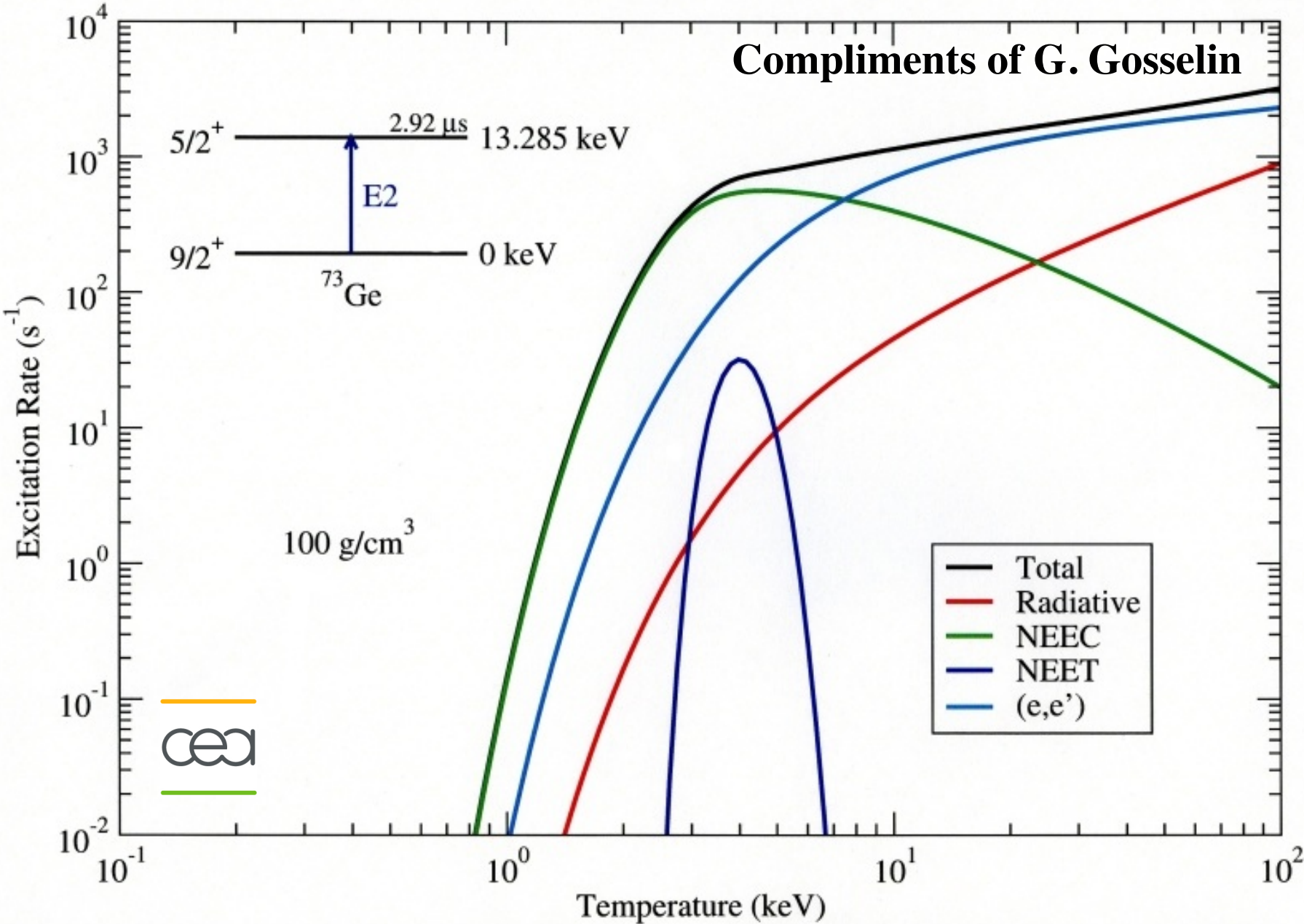


Atomic-nuclear (electron) interactions
NEEC, NEET, IES*
Time Reverse: IC-decay

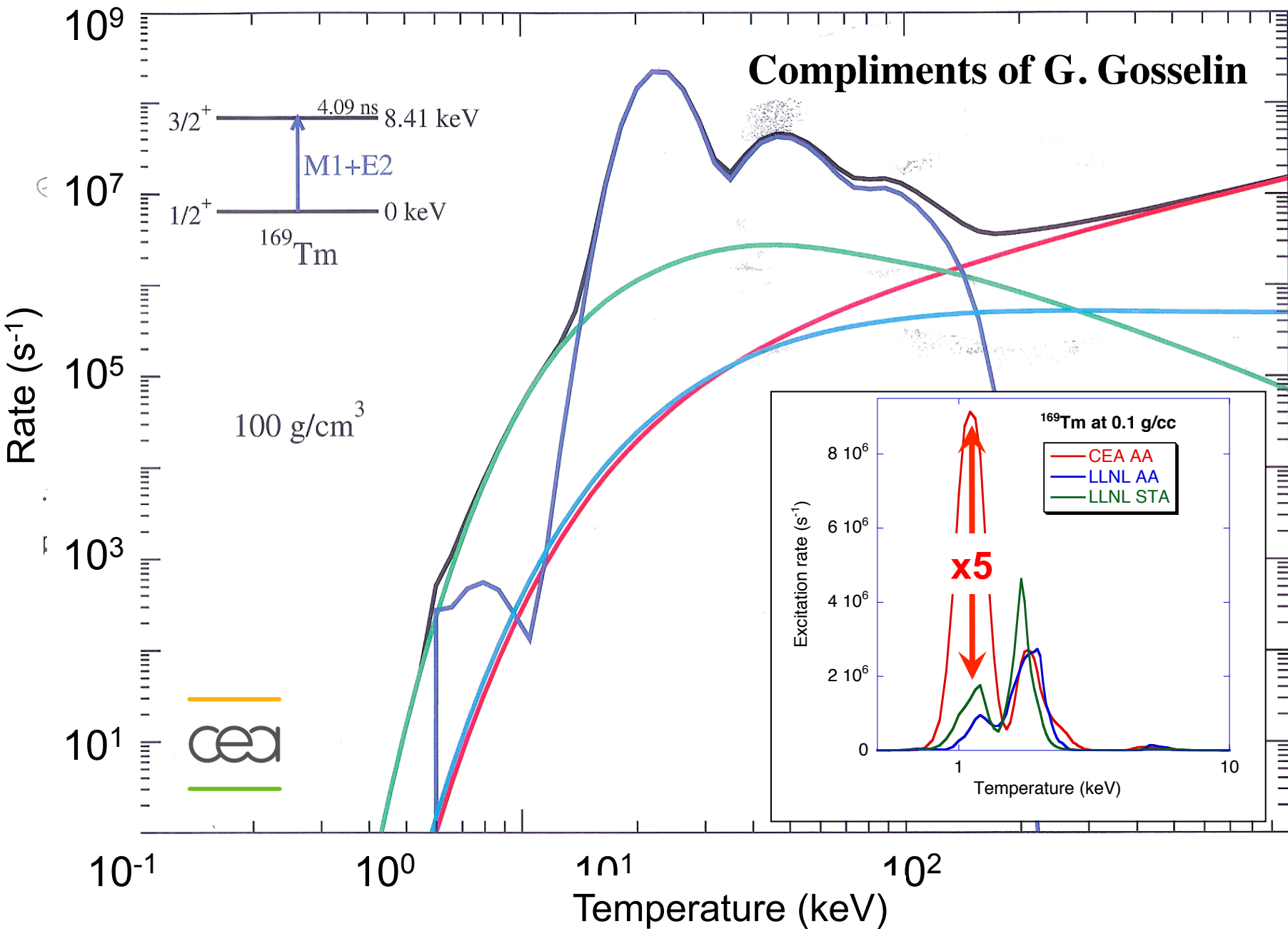


How rapid are these interactions?

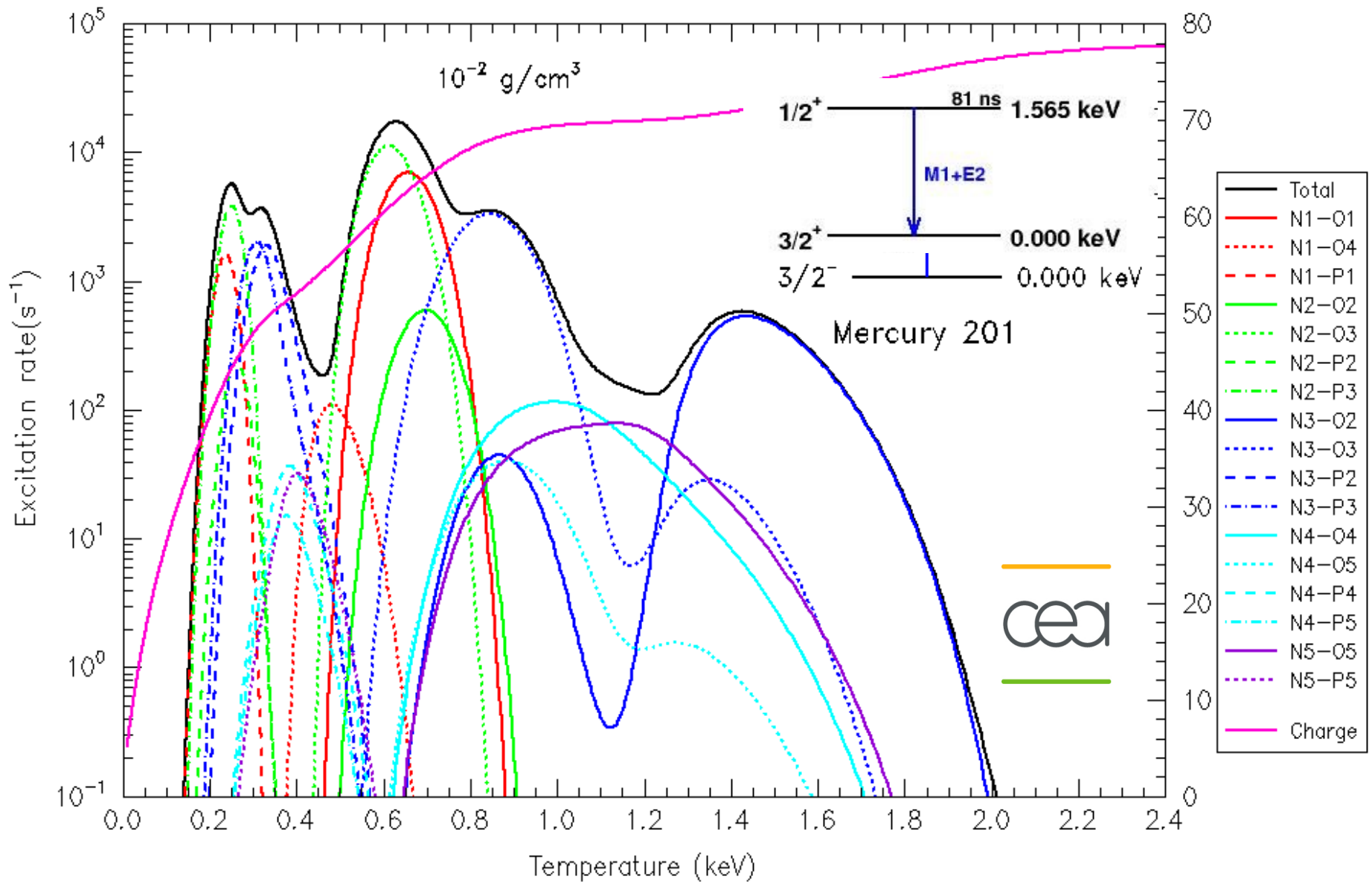
Excitation time scales for ^{73}Ge



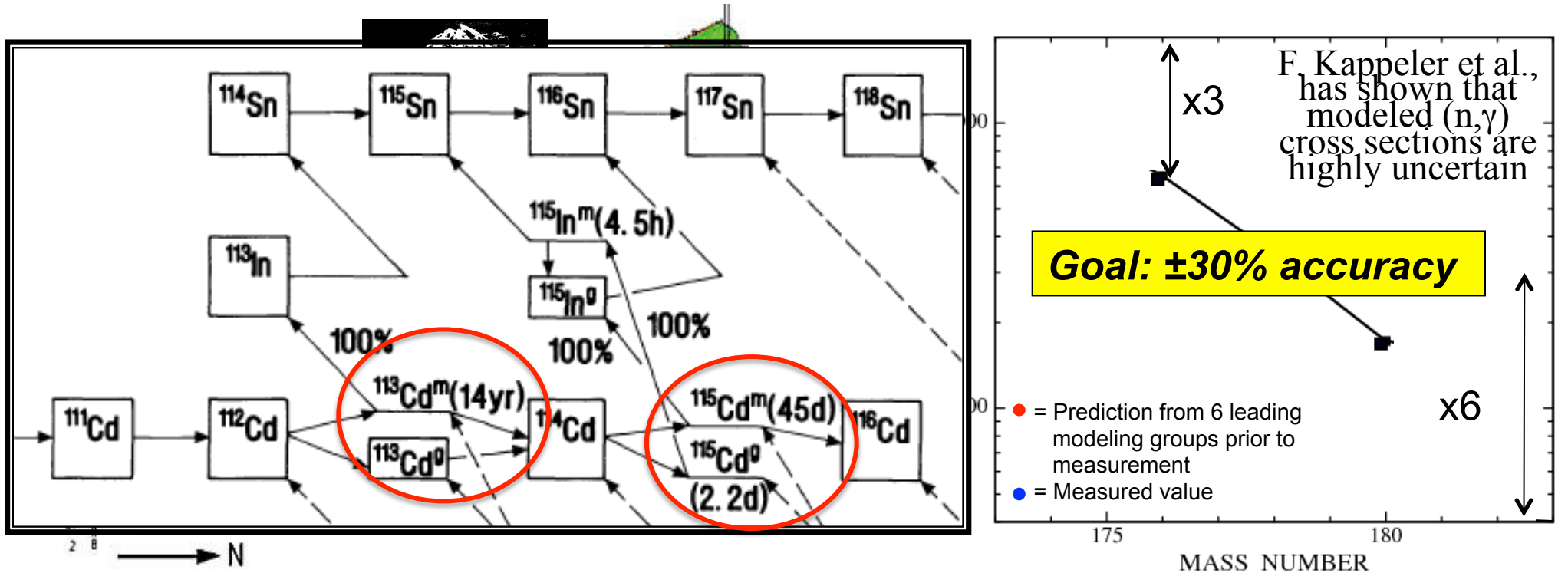
Excitation time scales for ^{169}Tm



NEEC/NEET rates are very sensitive to the underlying atomic structure



Roughly half of the elements with $26 \leq Z \leq 83$ are formed via slow neutron capture in an *astrophysical high energy density plasmas*



NIF @ 10^{14} neutrons crams 2800 years* of neutron capture into every shot

Can we use NIF to study the effects of the HEDP on (n,γ) nucleosynthesis?

*Busso, Gallino and Wasserburg, Annu. Rev. Astron. Astrophys. 1999. 37:239–309

R.A. Ward, Ap. J. **216**: 540-547, 1977, *Z.S. Nemeth et al., Ap. J. 426 357-365, (1994)*

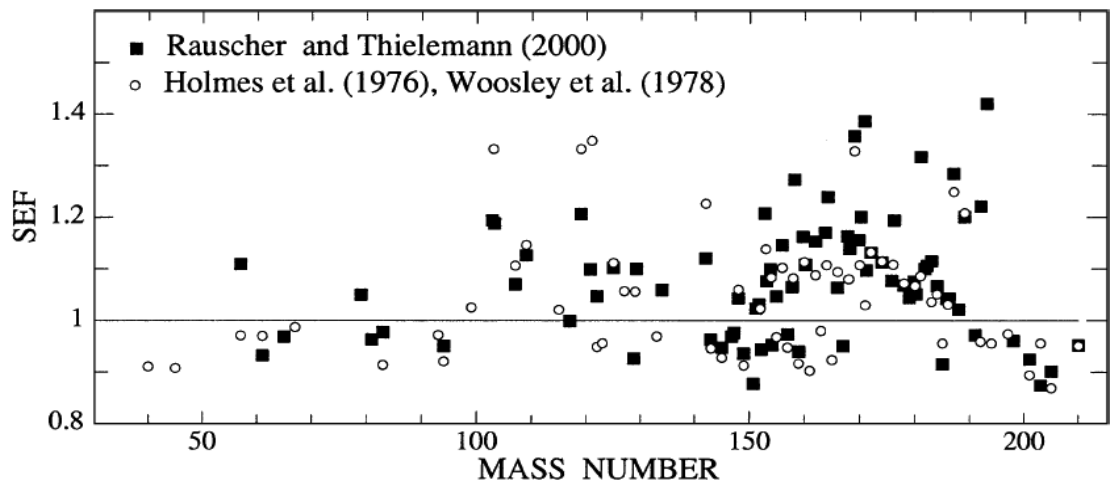
T. Hayakawa, et al., AIP Conf. Proc. **1238**, 225 (2010), doi: 10.1063/1.3455935

An additional complication is that many important* s-process branch point nuclei have low-lying excited states whose population can influence $\sigma_{(n,\gamma)}$ and β -decay lifetimes

S-process (n, γ) enhancement due to excited states*

$$SEF(kT) = \frac{\sigma_{HEDP}}{\sigma_{GS}} = \frac{\sum_{i=0}^{\infty} (2J_i + 1) \sigma(E_x = E_i) e^{-E_i/kT}}{\sigma_{GS} \sum_{i=0}^{\infty} (2J_i + 1) e^{-E_i/kT}}$$

*Bao & Kappeler At. Dat. Nucl. Dat. Tables **76**, 70–154 (2000)



Branch Point	Gnd State J^π	1 st Exc. State E_x (keV)	1 st Exc. State J^π
⁷⁹Se	7/2⁺	95.77	1/2⁻
⁸⁵Kr	9/2⁺	304.871	1/2⁻
¹⁴⁷Pm	7/2⁺	91.1	5/2⁺
¹⁵¹Sm	5/2⁻	4.821	3/2⁻
¹⁶³Ho	7/2⁻	100.03	9/2⁻
¹⁷⁰Tm	1⁻	38.7139	2⁻
¹⁷¹Tm	1/2⁺	5.0361	3/2⁺
¹⁷⁹Ta	7/2⁺	30.7	9/2⁺
²⁰⁴ Tl	2 ⁻	414.1	4 ⁻
²⁰⁵ Pb	5/2 ⁻	703.3	7/2 ⁻
¹⁸⁵W	3/2⁻	23.547	1/2⁻

NIF (or LMJ) are the *only* places where (n, γ) might be measured on ground+excited states

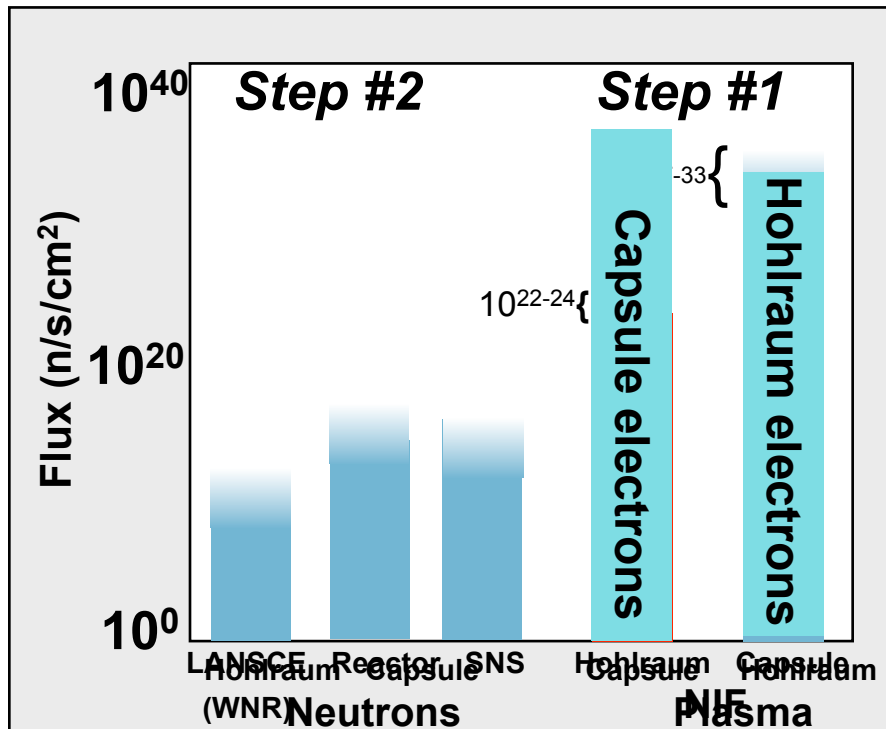
NIF concentrates all 192 laser beam energy
in a football stadium-sized facility into a mm^3

Matter
Temperature $>10^8$ K
Radiation
Temperature $>3.5 \times 10^6$ K
Densities $>10^3$ g/cm^3
Pressures $>10^{11}$ atm

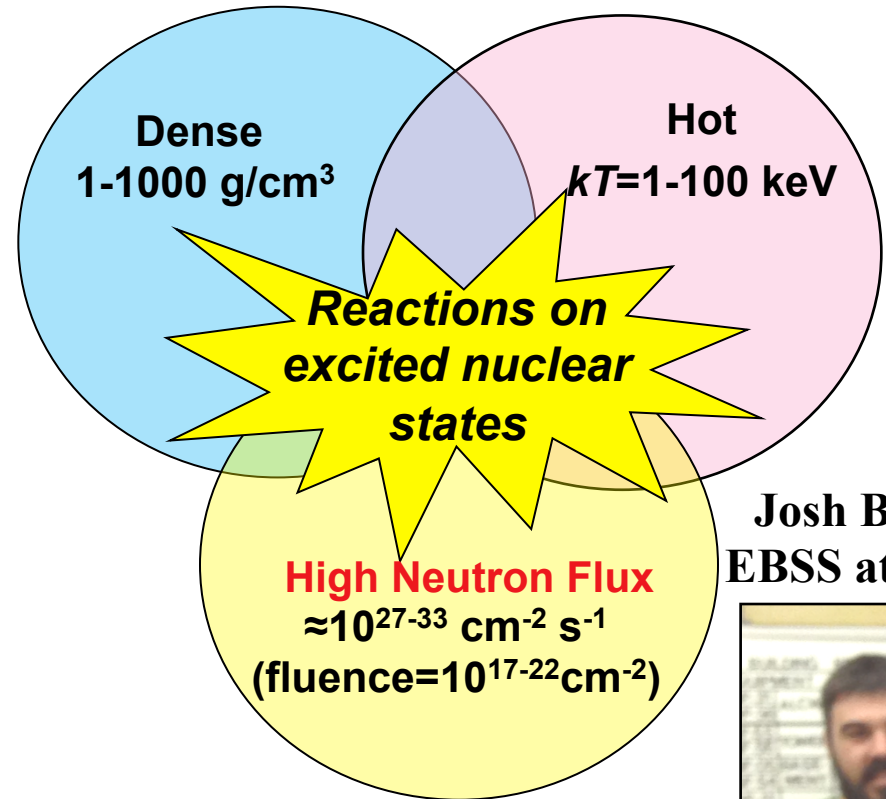




The high e , γ and n -flux in a NIF capsule might allow us to explore reactions on short-lived nuclear states



NIF capsule/hohlraum



Josh Brown
EBSS attendee

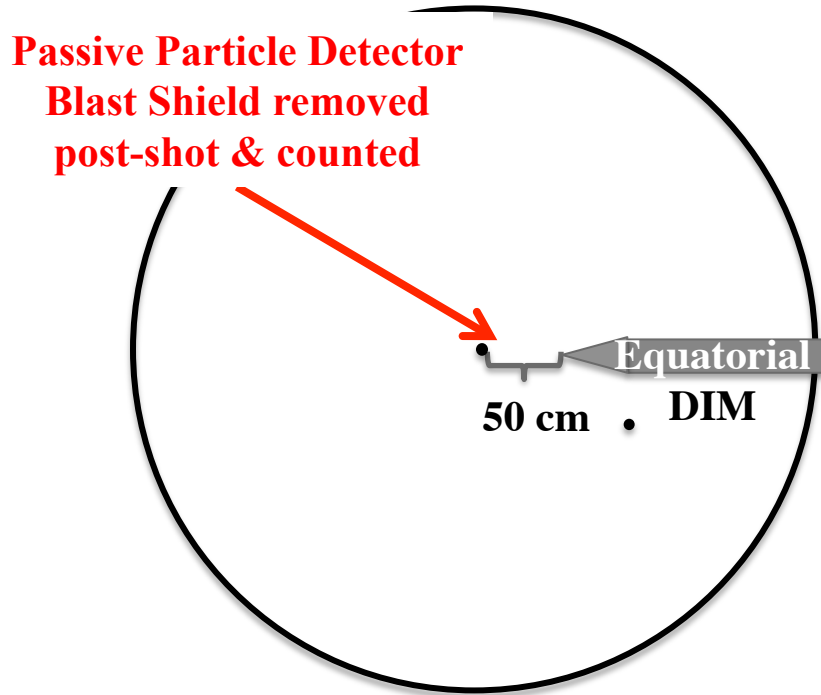


Excited State Reaction Possibilities

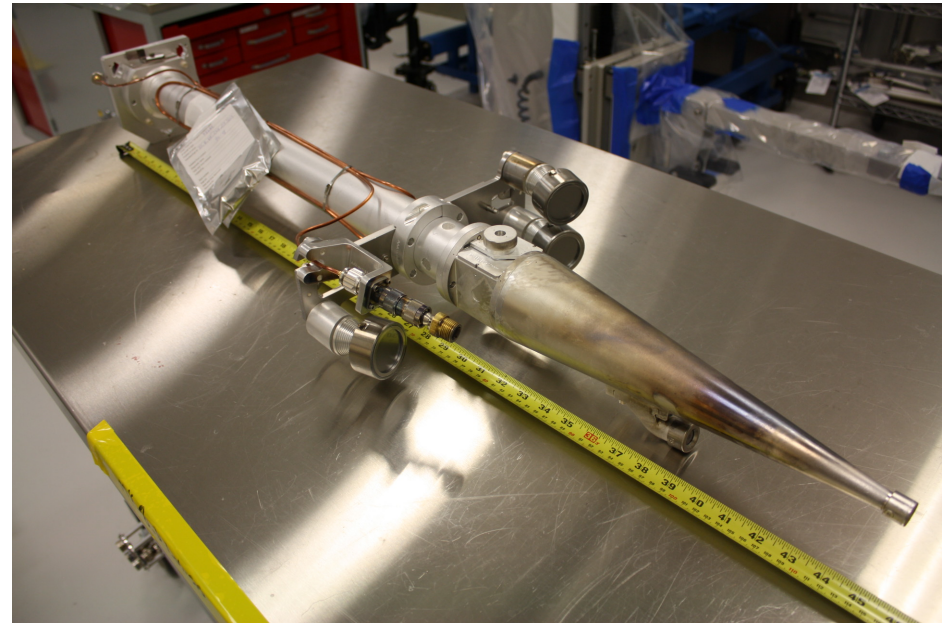
Option #1: Excite a target nucleus with the plasma then hit it with neutrons

Option #2: Excite a target nucleus with neutrons then interact with the plasma

First hints of NPI at NIF: Radioactive ^{196}Au and ^{198}Au from (n,2n) and (n, γ) on the ^{197}Au hohlraum



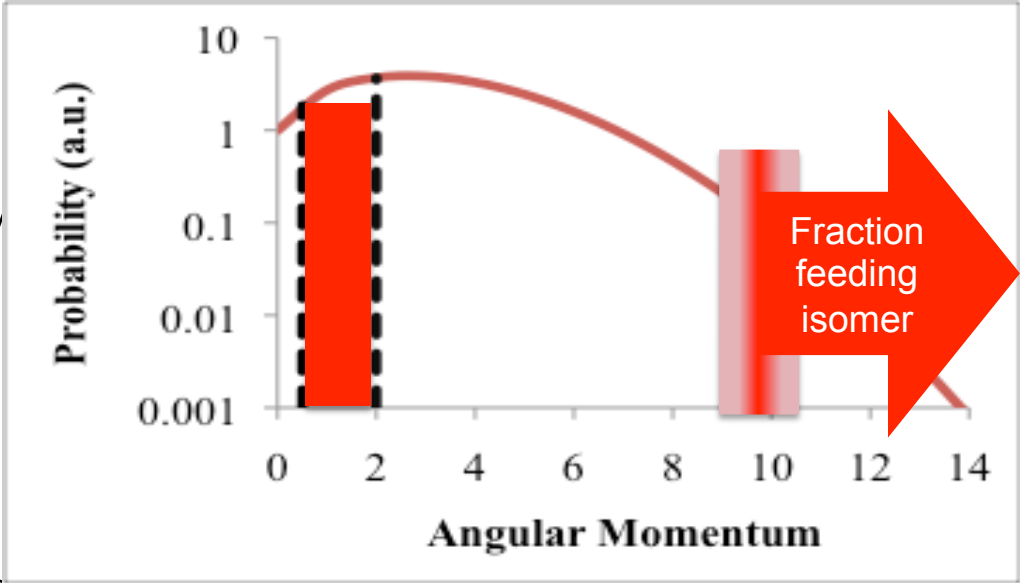
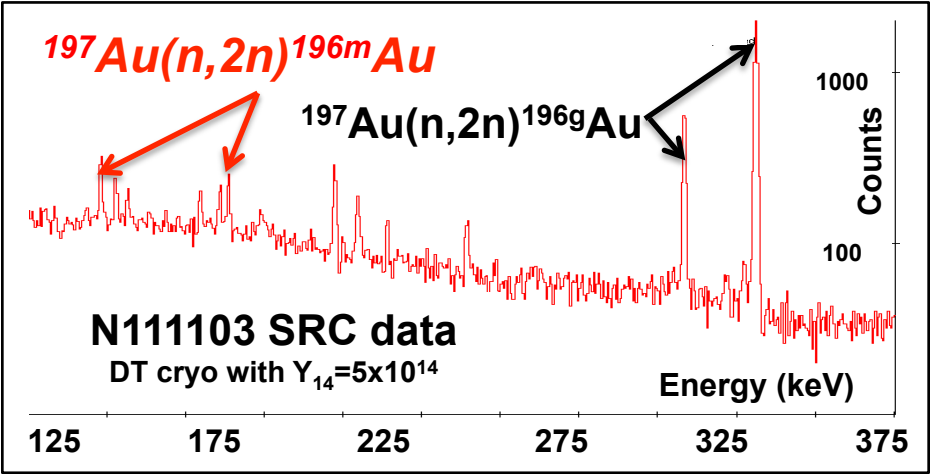
Diagnostic Insertion Manipulator (DIM)



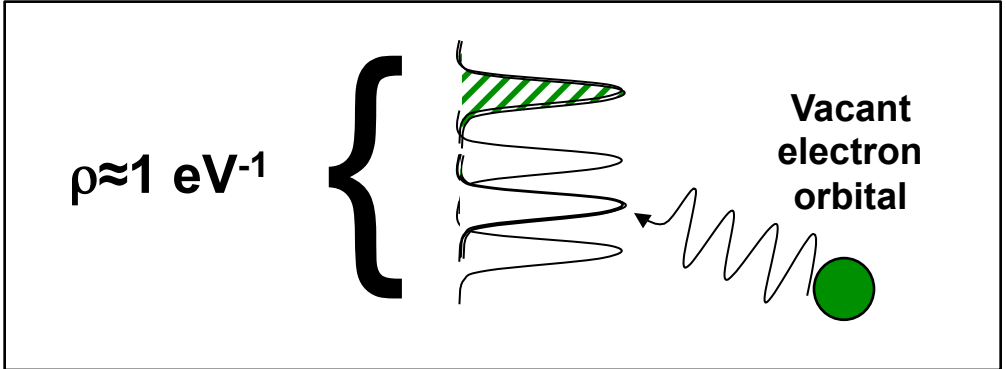
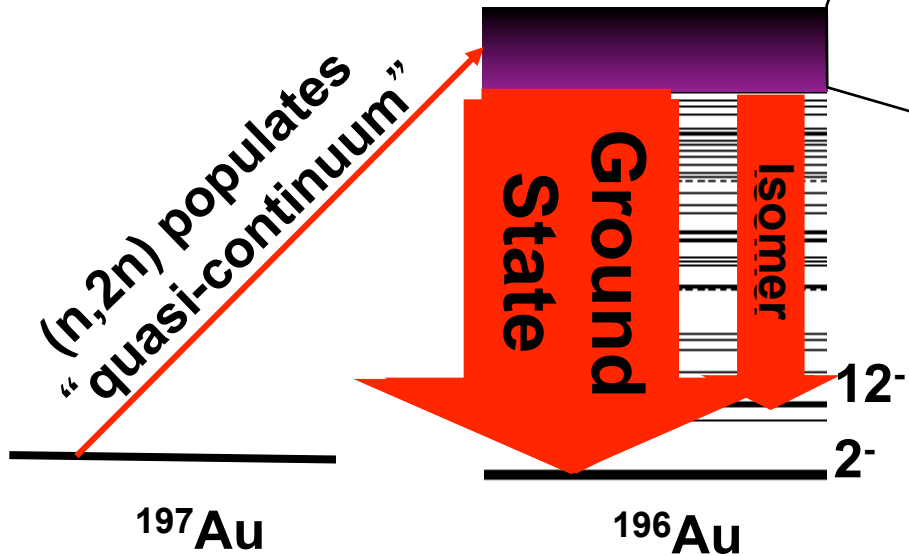
Time Sequence

1. Shot
2. 6-12 hours later DIM removed, samples collected and transported to Building 151 counting facility
3. 2-3 days later data becomes available

The 9.7 hour 12^- isomer in ^{196}Au might allow us to explore the interaction of highly-excited states with a HEDP?

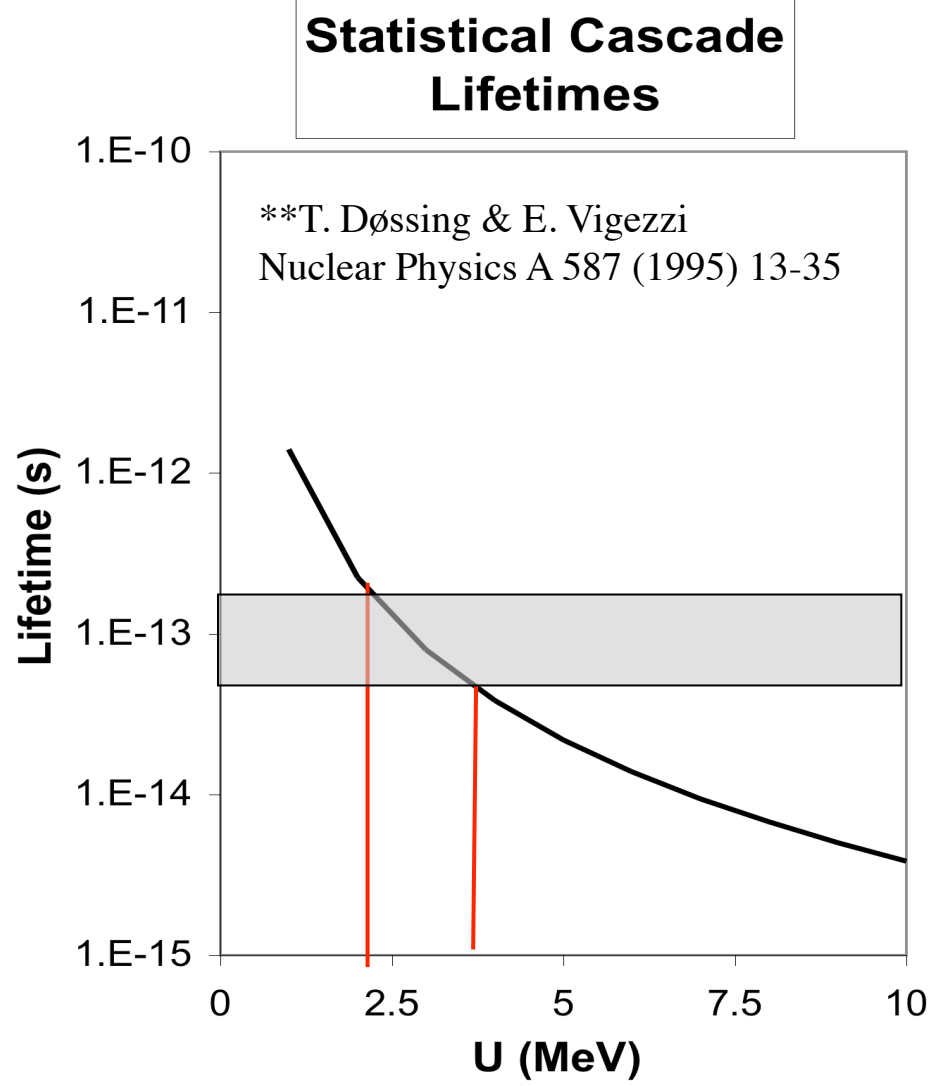
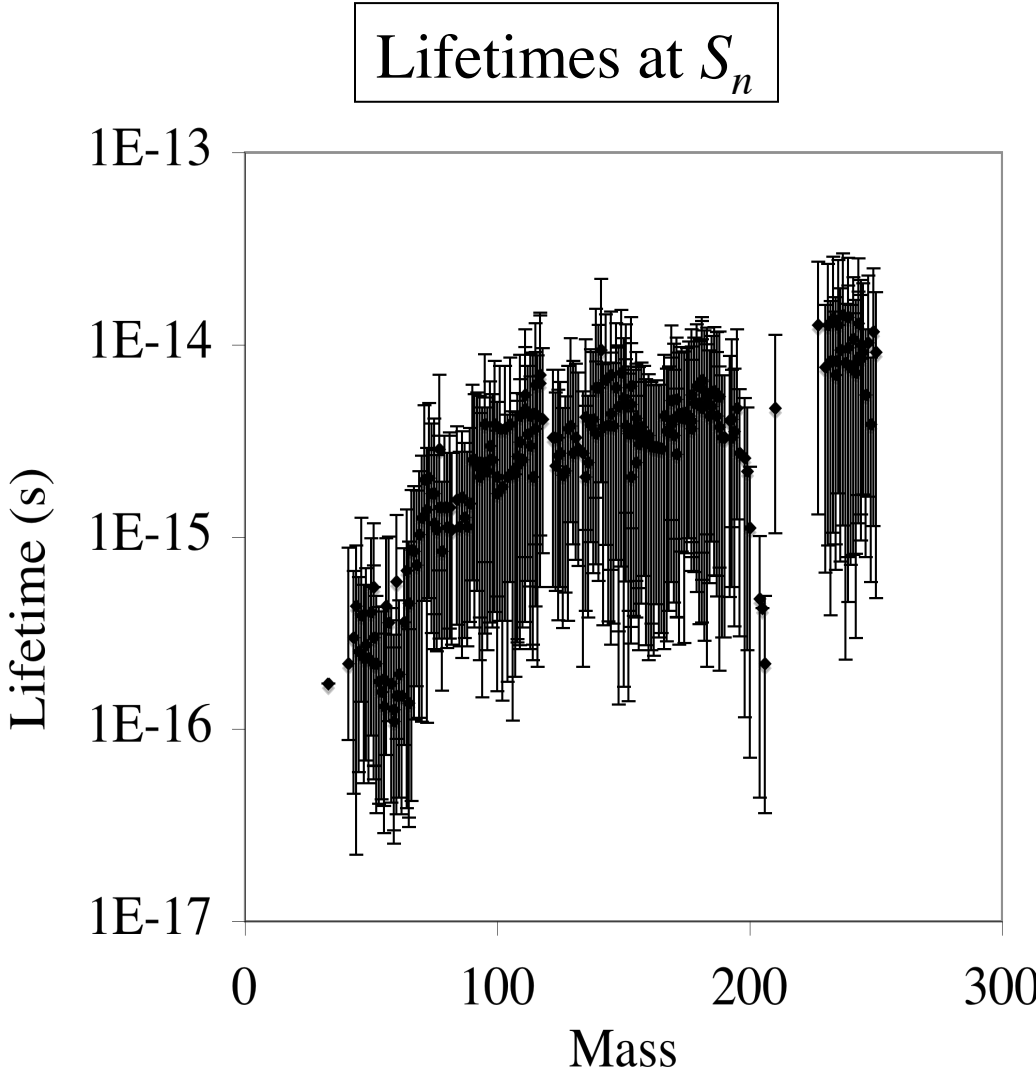


$S_n = 6.642 \text{ MeV}$



This is entirely new Nuclear Physics

A survey of (n,γ) resonance widths* shows that $E_x \approx 4-5 \text{ MeV}$ quasi-continuum lifetimes are on the order of $\tau_{DT-burn}/P$

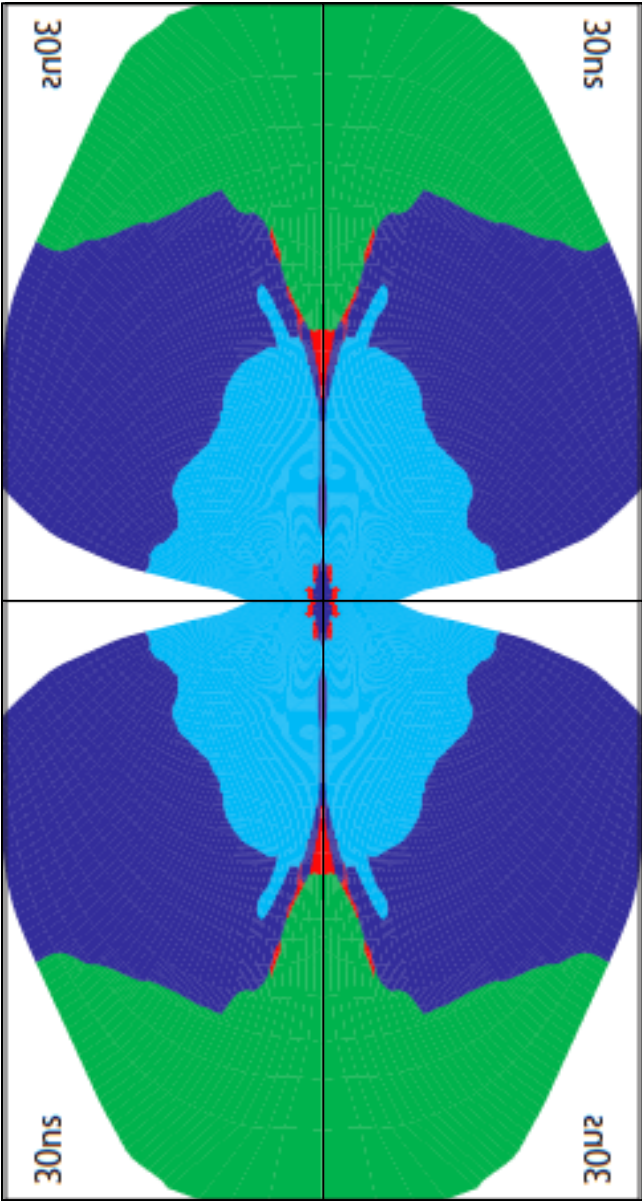
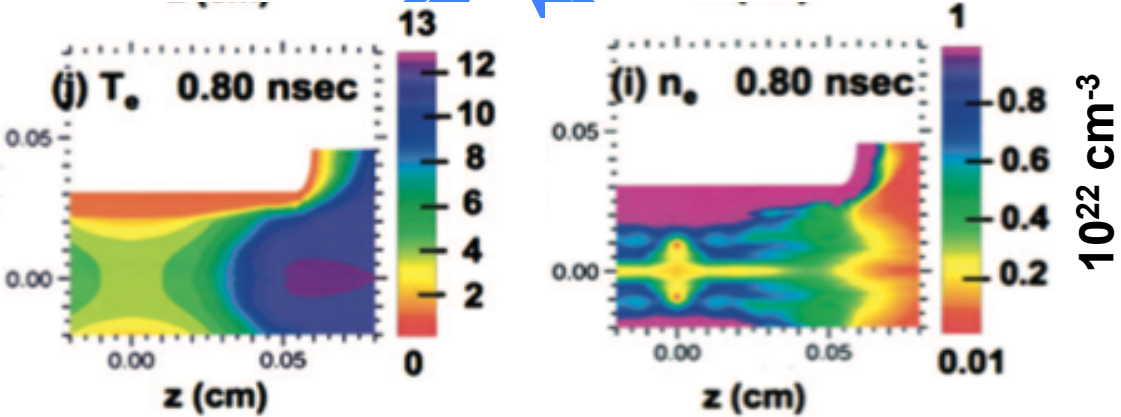
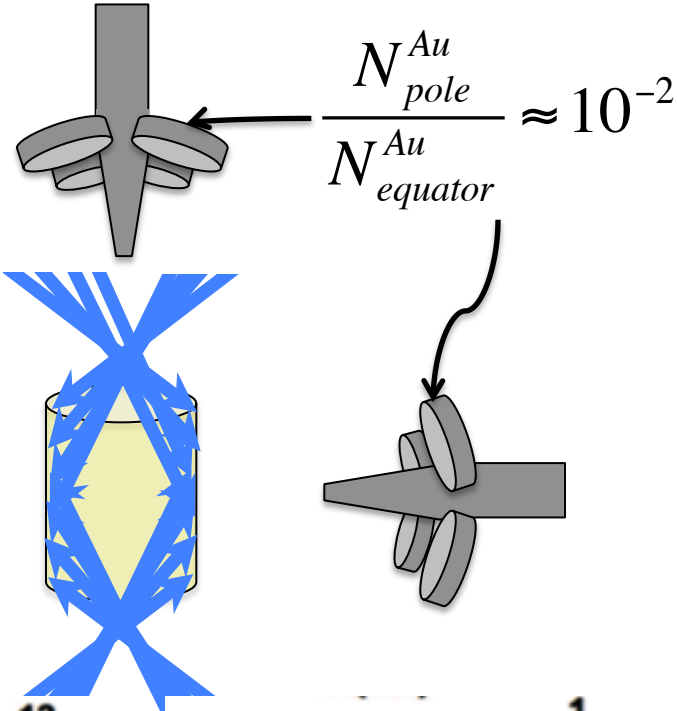


This could play a role in high-flux astrophysical scenarios (supernovae, etc.)

*RIPL-2 "obninsk" compilation

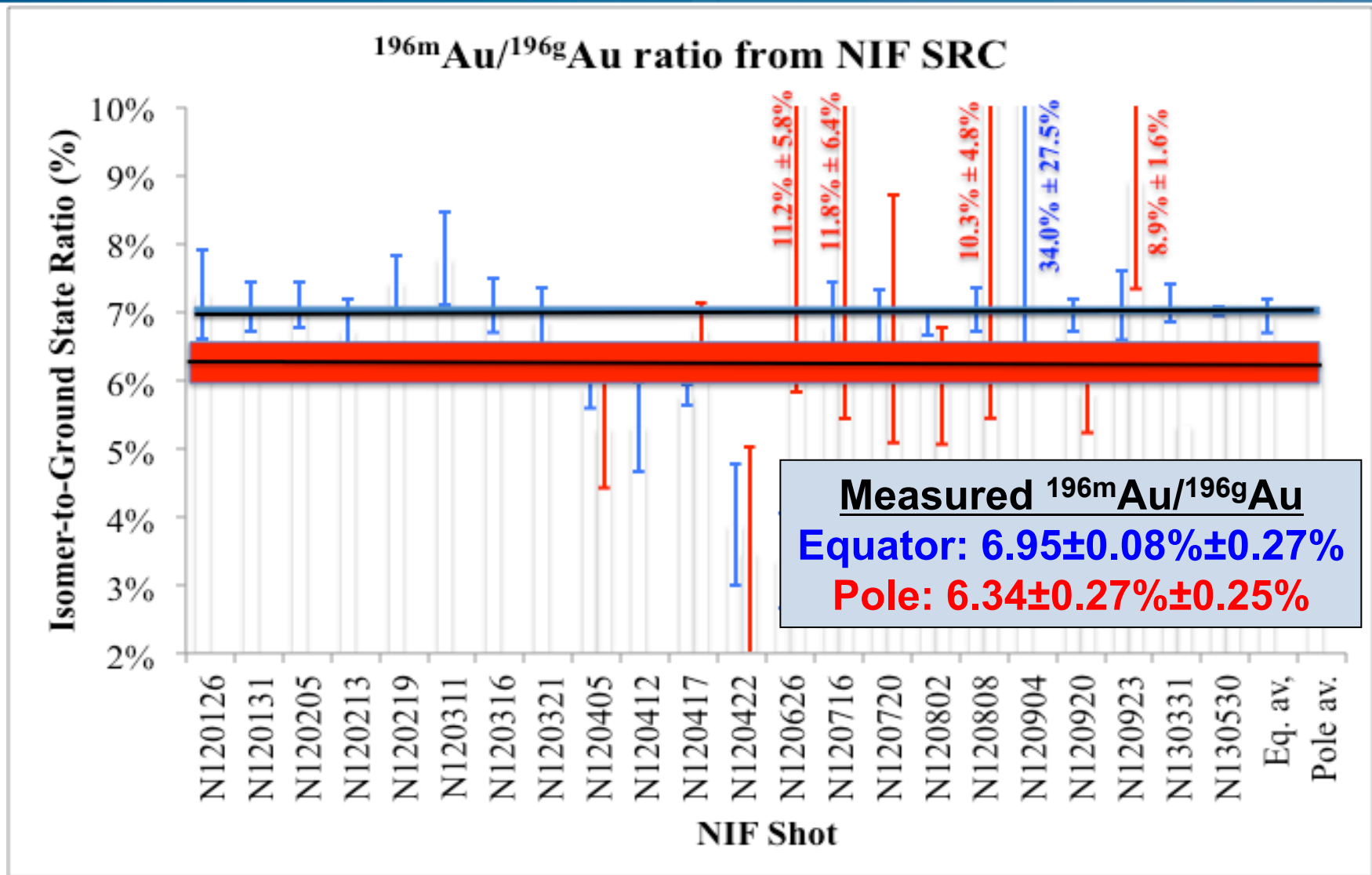
Radioactive ^{196}Au collected from the pole and waist of the NIF come from very different plasma conditions

D. Eder et al., UCRL-JRNL-206693



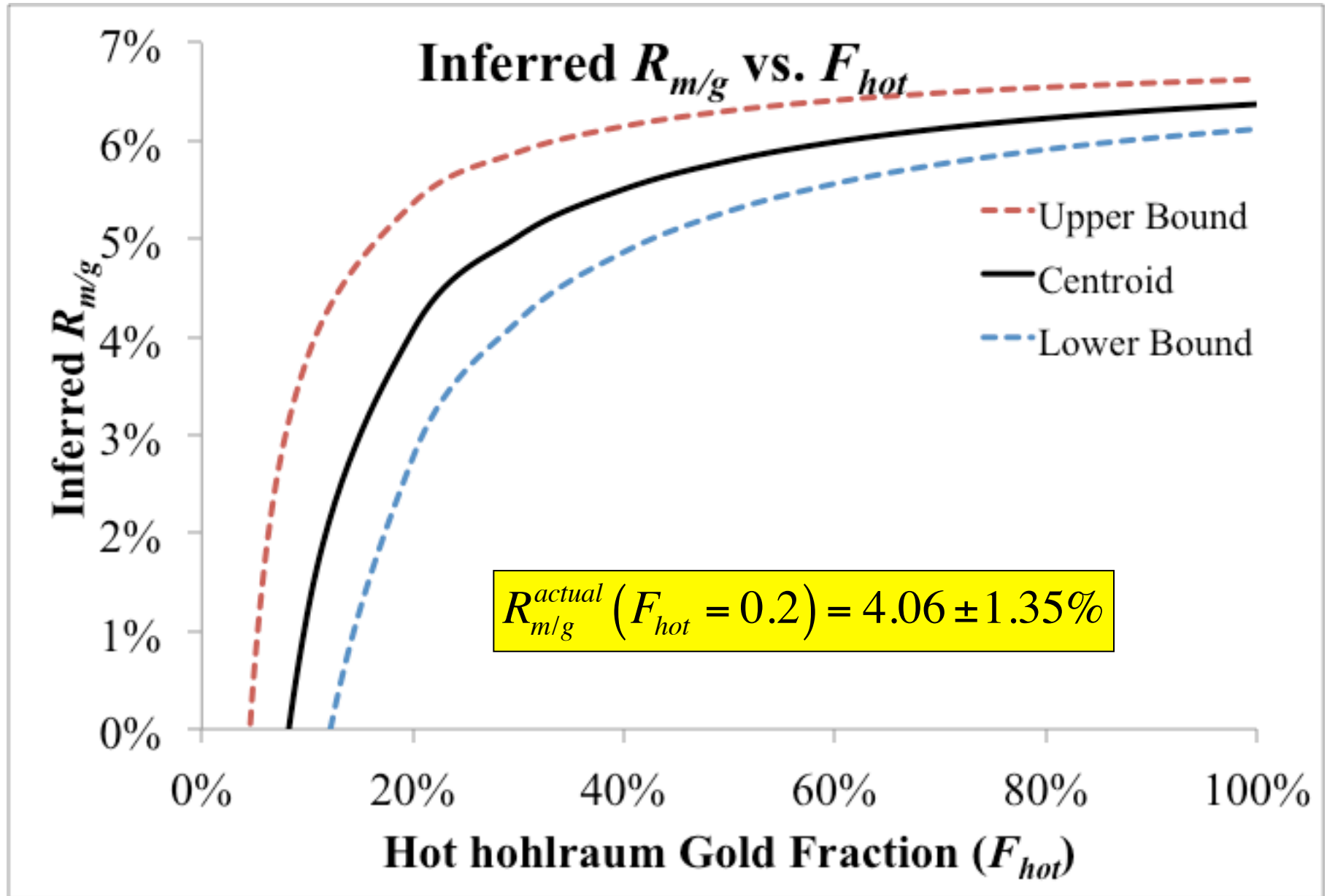
Polar Au comes from a HEDP while equatorial Au does not

Is debris from the NIF hohlraum suggesting that the $J^\pi=12^-$ isomer feeding is being effected by NPIs?



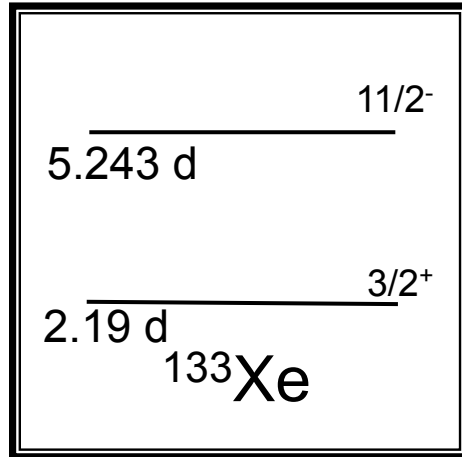
$$\frac{R_{m/g}^{pole}}{R_{m/g}^{equator}} = 0.91 \pm 0.04\%$$

If we assume a given fraction of the polar Au is hot we can determine the actual m/g ratio

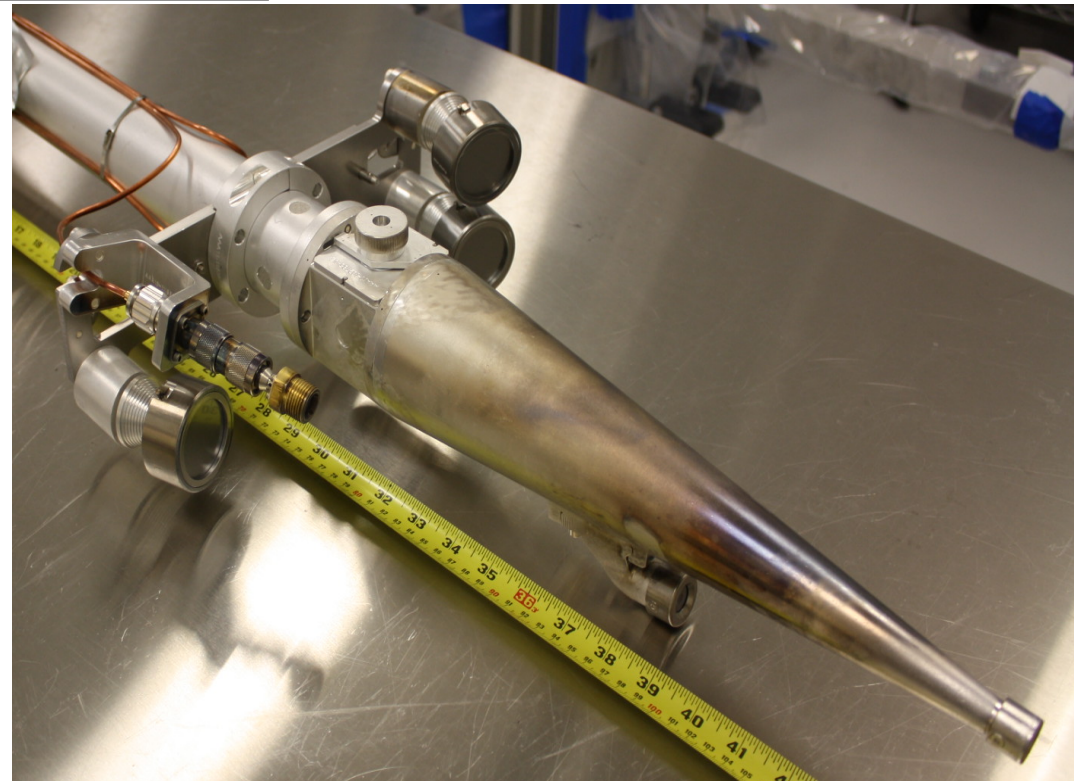
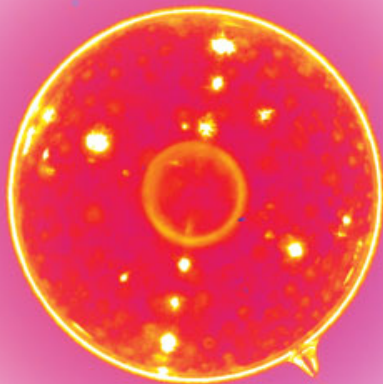


Option #2: A “better” NIF experiment using a ^{134}Xe -doped “exploding pusher” capsule

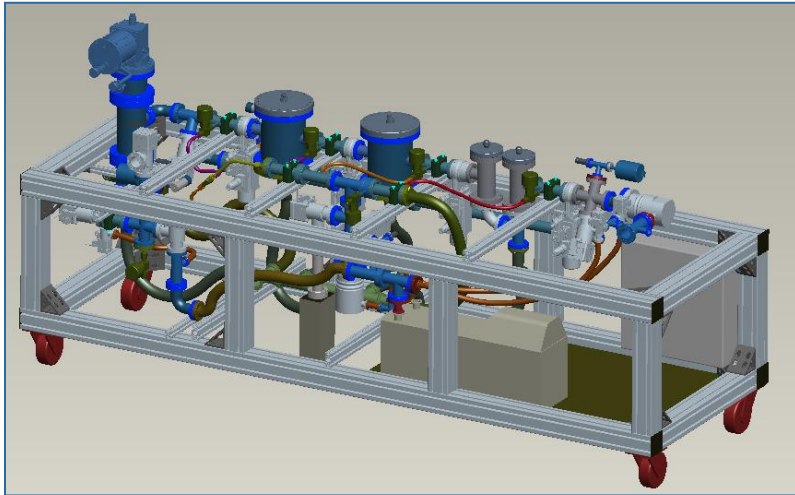
We maximize both neutron flux and plasma density by placing a ^{134}Xe dopant nuclei in a **direct-drive** target



...plus a “control” sample outside the plasma in a sample positioner 50cm from the target



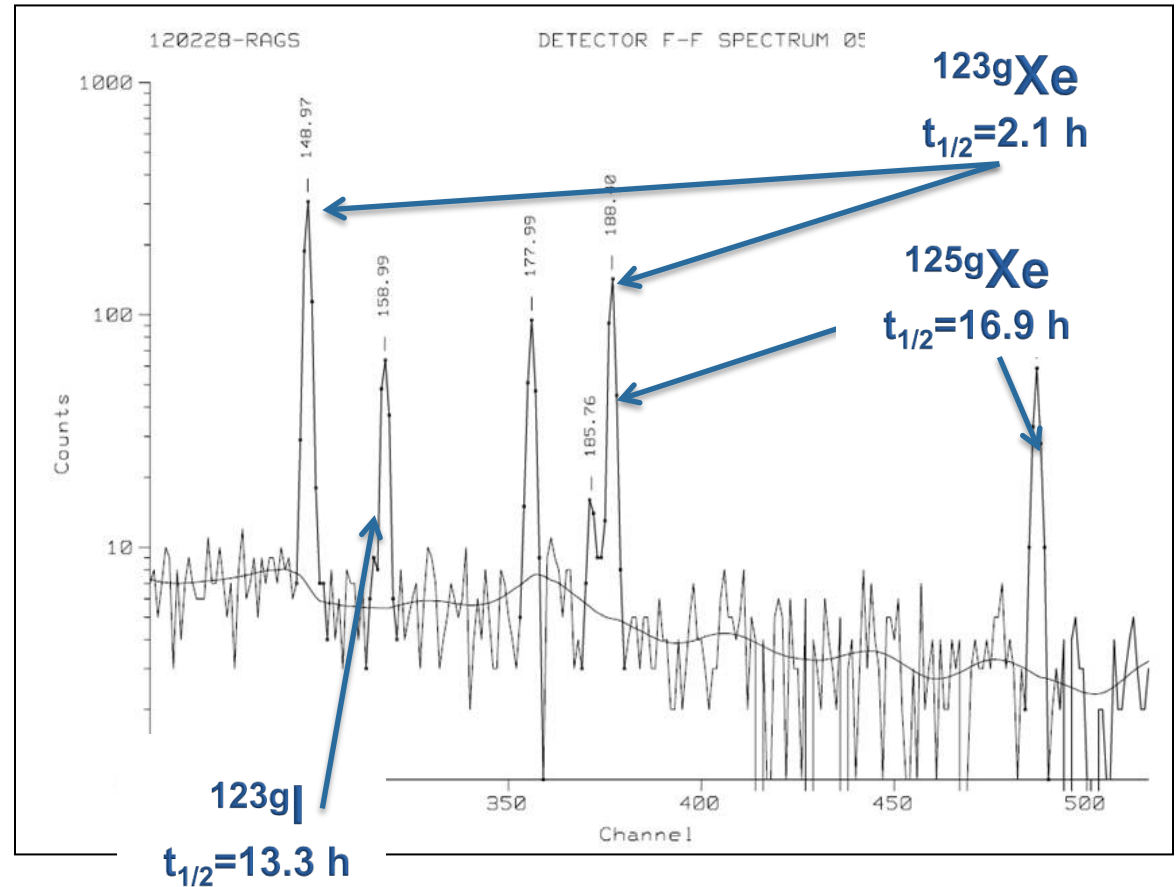
Radioactive $^{133m,g}\text{Xe}$ can be pumped out of NIF minutes after a shot using the RAGS (Radiochemical Analysis of Gaseous Samples) system



Exploding pusher test: $^{124}\text{Xe}, ^{126}\text{Xe}$ -doped capsule
NIF shot N120228-001-999

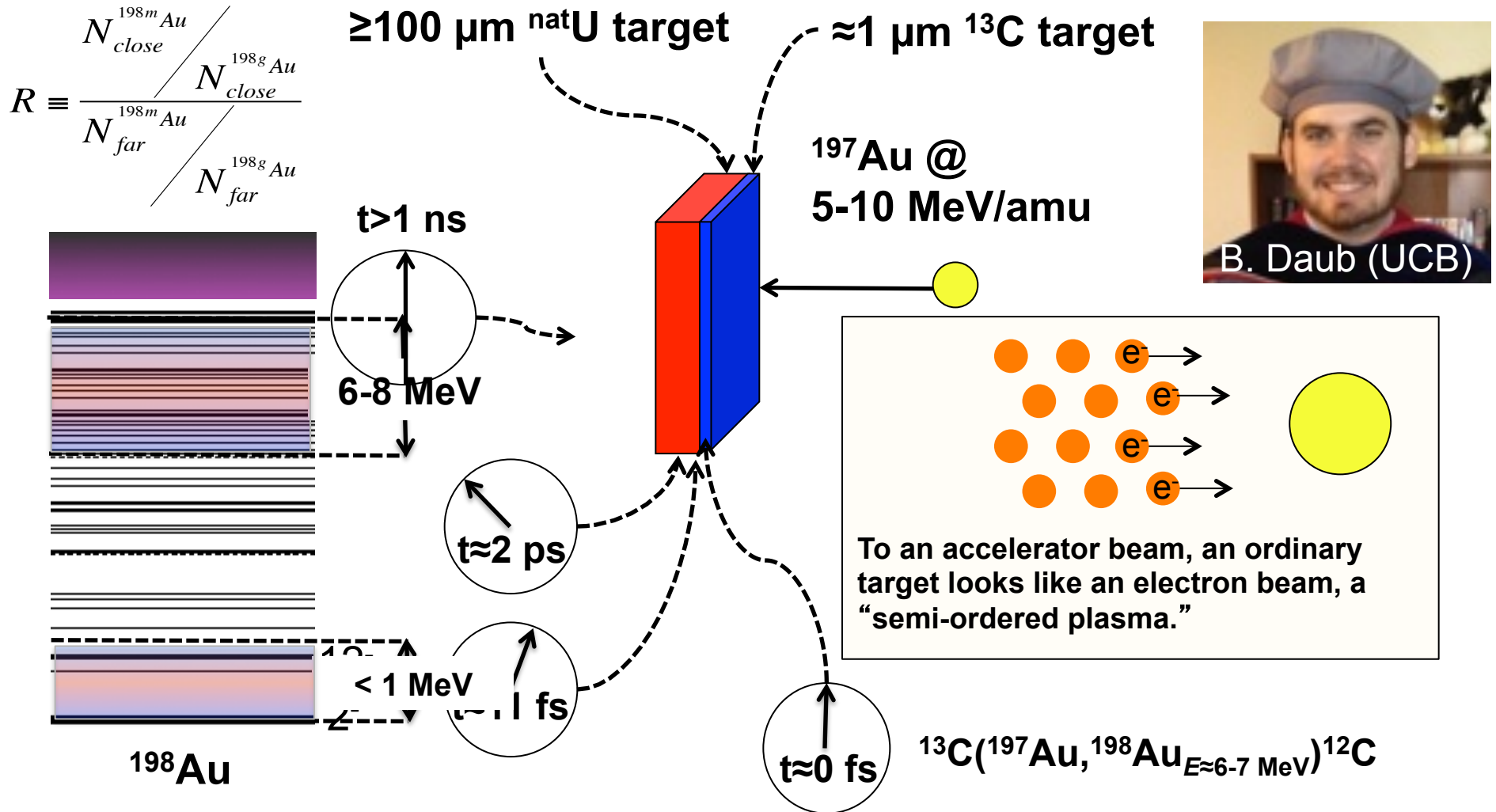
NPI effects can be observed using the *Double-Isomer-to-Ground State (DIGS) Ratio*

$$R_{DIGS} \equiv \frac{\frac{N_{capsule}^{133m}\text{Xe}}{N_{capsule}^{133g}\text{Xe}}}{\frac{N_{SRC}^{133m}\text{Xe}}{N_{SRC}^{133g}\text{Xe}}}$$



Collection efficiency > 63% has been demonstrated

Option #3: A complementary accelerator-based experiment can also be performed using *GeV* Au beams

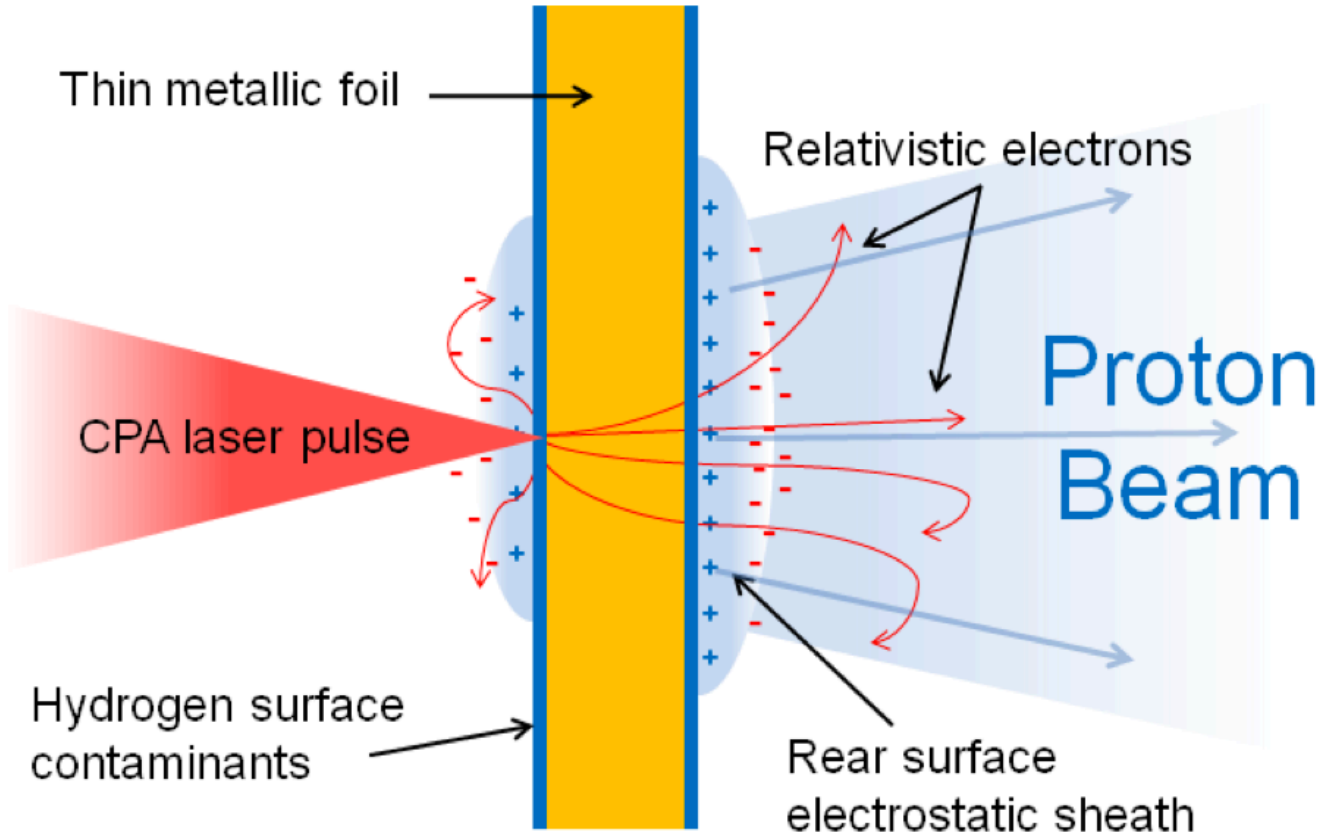


First test experiment fielded at LBNL 3/13
 – ¹⁹⁸Au formed, but no isomer was formed
 due to low beam energy (4.2 MeV/amu)

Plasma Properties	NIF	LBNL
Electron Fluence (cm ⁻²)	≈3x10 ²²	≈10 ²⁰
Temperatures (keV)	T _e ≈5-50, T _g =0.3	T _e ≈2-20, T _g =n.a.

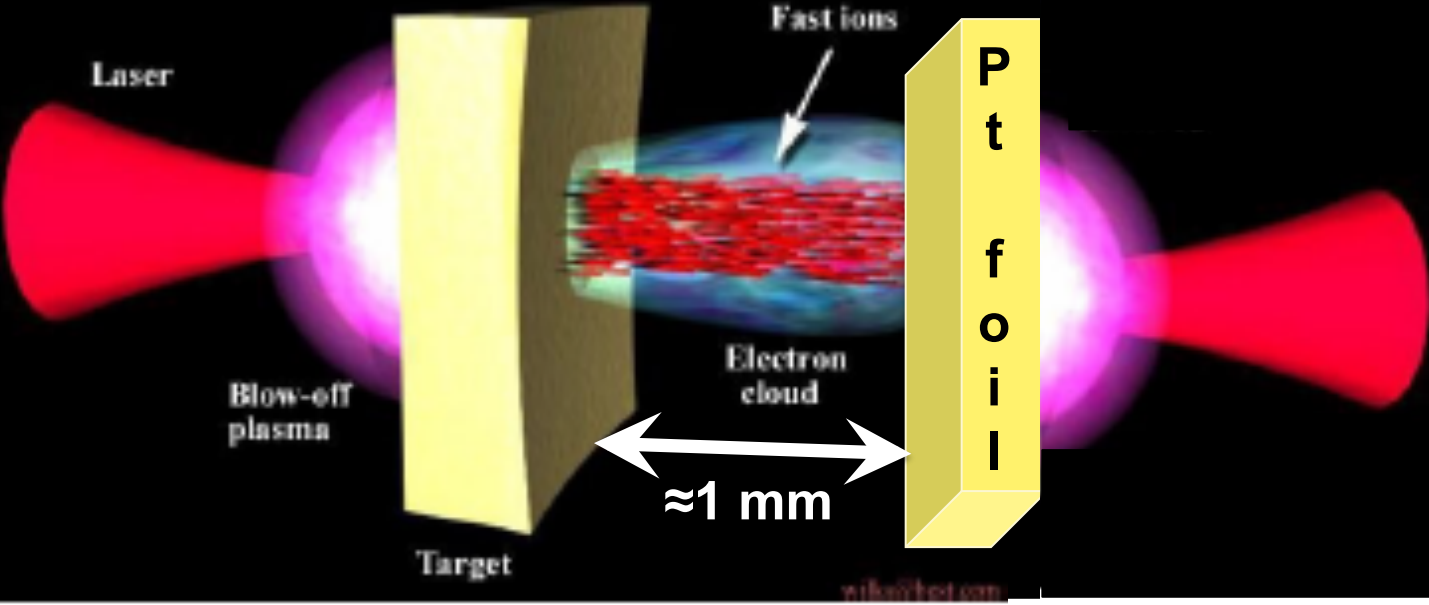
Option #4: We can use protons from a petawatt laser to make excited ^{196}Au via $^{198}\text{Pt}(p,3n)$

Target Normal Sheath Acceleration



TNSA proton based nuclear-plasma experiment make $^{196m,g}\text{Au}$ using the $^{198}\text{Pt}(p,3n)$ reaction

Use TNSA protons from a petawatt laser to make an excited nucleus via the $^{198}\text{Pt}(p,3n)^{196m,g}\text{Au}$



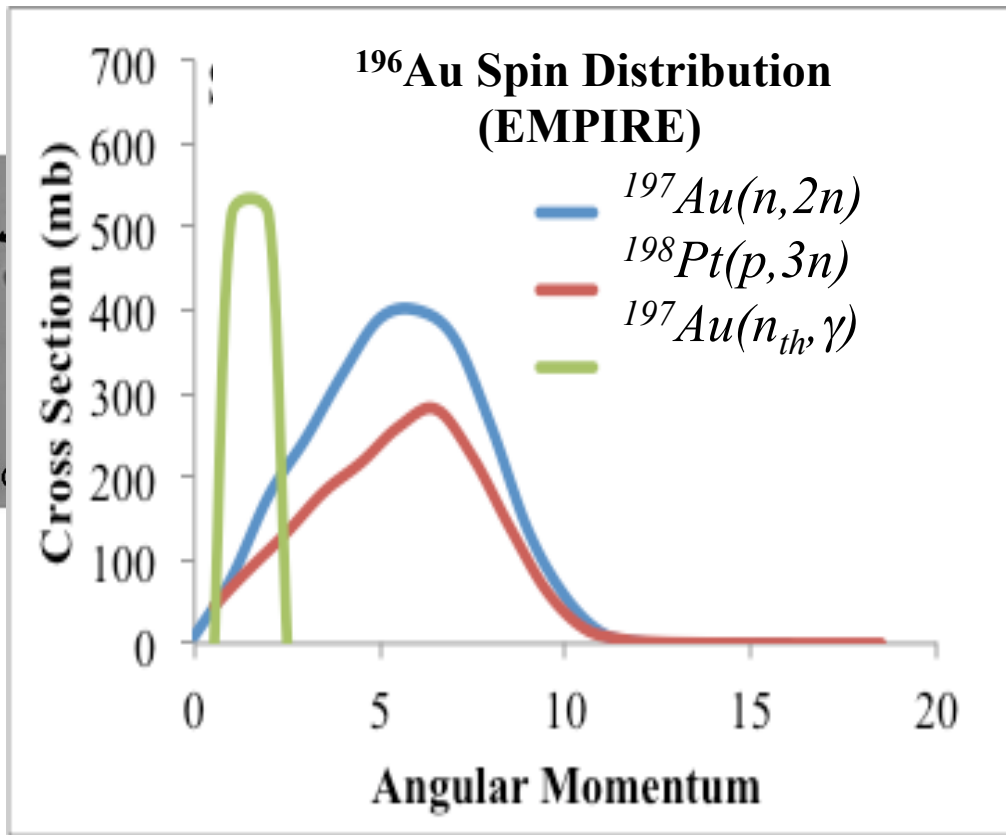
Use a long pulse (ns) laser to place the target nuclides into an HED plasma state a fixed amount of time after the TNSA protons arrive

First experiment: Platinum in a plasma state when the protons hit

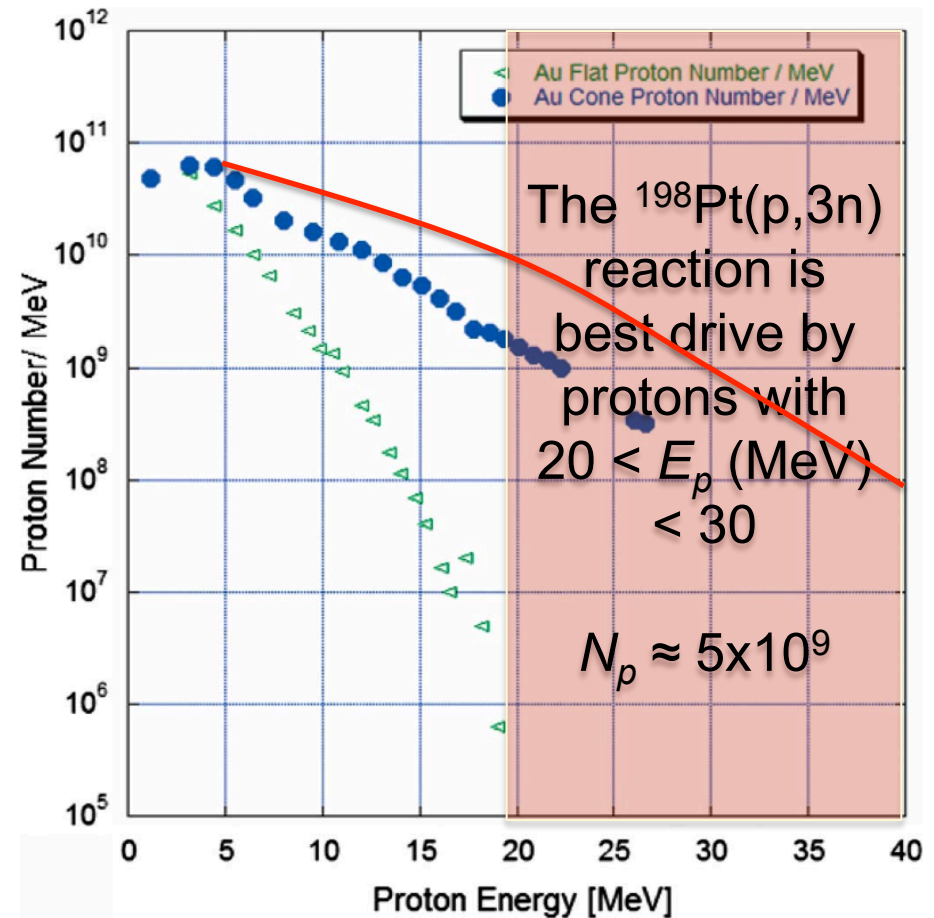
Control experiment: Platinum put into a plasma state *after* the protons hit

The TNSA proton spectrum can be estimated using recent “state of the art” results

- Results from Flippo (2008) at LANL show >10-fold increase in high-energy proton production in shaped targets *Laser power < 100 TW*



(a) \longleftarrow $\xrightarrow{200\mu\text{m}}$

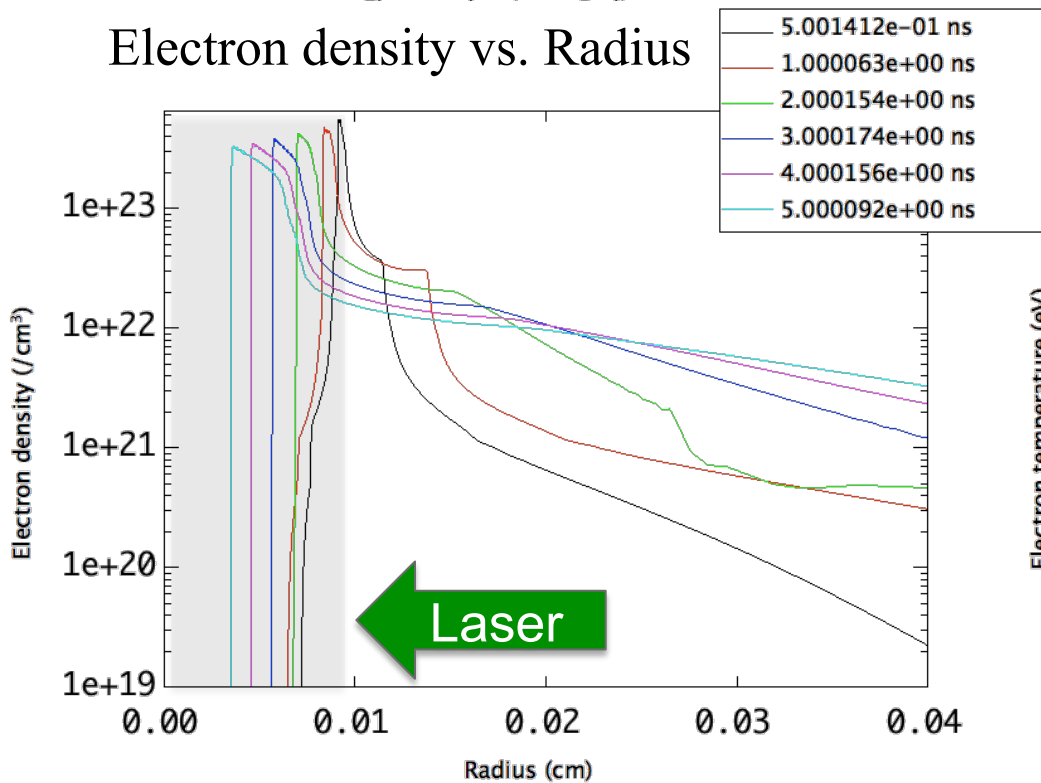


More recent results from *Roth* suggest an even harder proton flux
 From the related **BOA** mechanism (**B**reak **O**ut **A**fterbuner)

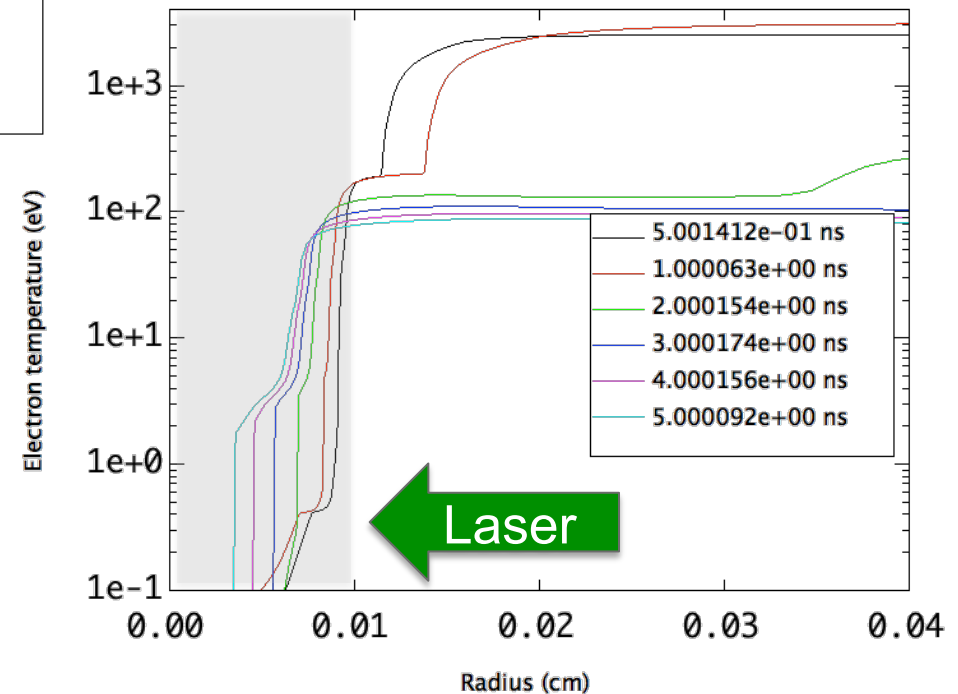
Long-pulse laser produces a variety of plasma conditions

1D Radiation Hydrodynamics simulations complements of P.F. Davis

Electron density vs. Radius



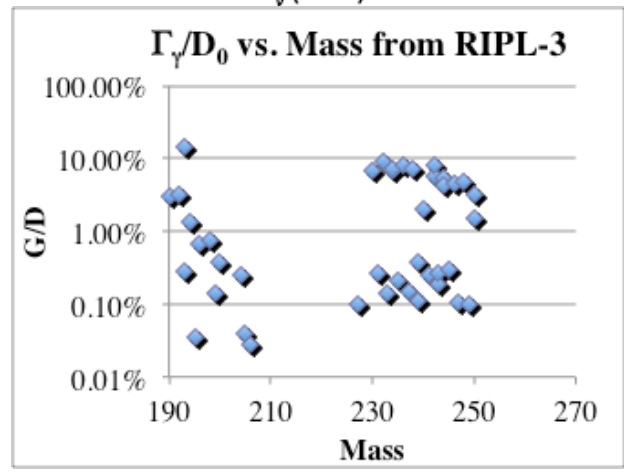
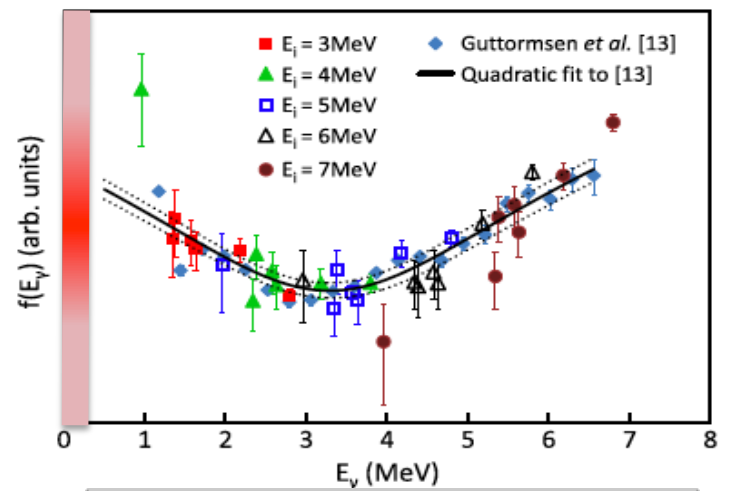
Electron temperature vs. Radius



Plasma Properties	NIF	TNSA
Electron Fluence (cm ⁻²)	≈3x10 ²²	≈10 ²⁰⁻²¹
Temperatures (keV)	T _e ≈5-50, T _g =0.3	T _e ≈0.2-3, T _g =0.2

Summary

- Interactions between highly-excited nuclear states and HEDPS can profoundly effect nucleosynthesis
- We have hints of this happening right now at NIF
- Outstanding questions:
 - What are the appropriate atomic rates?



NIF

